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# Design of a High Performance Selective Solar Absorber with the Structure of SiO<sub>2</sub>-TiO<sub>2</sub>-TiN<sub>x</sub>O<sub>y</sub>-Cu

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Using TiN<sub>x</sub>O<sub>y</sub> thin film as the absorbing layer, a spectrally selective solar absorber (SSA) with the structure of SiO<sub>2</sub>-TiO<sub>2</sub>-TiN<sub>x</sub>O<sub>y</sub>-Cu has been designed; and the SSA performance has been evaluated by both calculation and experiment. The TiN<sub>x</sub>O<sub>y</sub> layer is fabricated by RF magnetron sputtering of a pure TiN target with Ar and N<sub>2</sub> gases. The introduction of oxygen in the TiN<sub>x</sub>O<sub>y</sub> is realized by using the residual oxygen in the deposition chamber. It is found that the N<sub>2</sub> flow rate has a large effect on the complex refractive index of the TiN<sub>x</sub>O<sub>y</sub> thin film and thus the SSA performance can be optimized by controlling the N<sub>2</sub> flow rate. With the TiN<sub>x</sub>O<sub>y</sub> absorbing layer deposited with the N<sub>2</sub> flow rate of 2 sccm, the SSA achieves the highest solar thermal conversion efficiency for a solar radiation concentration factor of equal to or larger than 5. The SSA based on the TiN<sub>x</sub>O<sub>y</sub> layer deposited with the N<sub>2</sub> flow rate of 2 sccm has a solar absorbance of 96.29% and a thermal emittance of 6.11% at the temperature of 400 °C.  
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Being the most abundant source of renewable energy, the solar energy resource exceeds all other renewable and fossil-based energy resources combined.<sup>1</sup> One of the direct ways to harvest the solar energy is to convert the solar energy into thermal energy, such as in the cases of solar water heater, solar thermo-photovoltaics, and concentrated solar power (CSP). CSP offers some unique advantages such as higher energy-conversion efficiency, higher thermal energy storage capability, and the potential to retrofit current coal power plants.<sup>2</sup> In a CSP system, the solar absorber has to maximally absorb solar energy while minimizing losses due to black body emission.<sup>2-4</sup>

Different types of selective solar absorber (SSA), such as semiconductor-metal tandems, metal-dielectric multi-layers and conductor-dielectric composites, etc. have been reported.<sup>5</sup> Due to its good thermal stability, high solar absorbance, and low infrared emittance, TiN<sub>x</sub>O<sub>y</sub> has been studied as solar absorbing layer.<sup>6-8</sup> Because the TiN<sub>x</sub>O<sub>y</sub> thin film can be treated as the composite of TiN and TiO<sub>2</sub>,<sup>9</sup> the conductive TiN serves as the absorption center while the TiO<sub>2</sub> is used as the dielectric host. This is similar to a cermet, such as W-Al<sub>2</sub>O<sub>3</sub>,<sup>10</sup> Cr-Cr<sub>2</sub>O<sub>3</sub>,<sup>11</sup> Mo-Si<sub>3</sub>N<sub>4</sub>,<sup>12</sup> in which the metal nanoparticles are dispersed in the dielectric matrix. For TiN<sub>x</sub>O<sub>y</sub> films, their optical properties can be tuned by adjusting the N/O ratio of the films.<sup>6,13,14</sup> For example, it was reported that optical properties of the TiN<sub>x</sub>O<sub>y</sub> films fabricated by reactive mid-frequency magnetron sputtering from titanium nitride target can be controlled in terms of the reactive gas flow rate.<sup>6</sup> This adds the flexibility to optimize the performance of solar absorber.

In this work, a high performance spectrally selective solar absorber (SSA) with the structure of SiO<sub>2</sub>-TiO<sub>2</sub>-TiN<sub>x</sub>O<sub>y</sub>-Cu has been designed by tuning the optical properties of the TiN<sub>x</sub>O<sub>y</sub> absorbing layer in terms of the N<sub>2</sub> flow rate used in the sputtering deposition of the TiN<sub>x</sub>O<sub>y</sub> layer. The TiN<sub>x</sub>O<sub>y</sub> absorbing layer is fabricated by RF magnetron sputtering of a pure TiN target with Ar and N<sub>2</sub> gases together with the residual oxygen in the deposition chamber. It is found that the optical properties in a wide spectral region from ultra-violet (UV) to infrared of the TiN<sub>x</sub>O<sub>y</sub> layer can be significantly changed by changing the N<sub>2</sub> flow rate as a result of the change in the chemical stoichiometry of the TiN<sub>x</sub>O<sub>y</sub> layer. It is shown that the refractive index and extinction coefficient of the TiN<sub>x</sub>O<sub>y</sub> layer can be easily tuned to maximize the SSA performance in the wide spectral region by varying the N<sub>2</sub> flow rate. This provides a simple but effective way for the design of a high performance SSA. With the TiN<sub>x</sub>O<sub>y</sub> absorbing layer deposited with the N<sub>2</sub> flow rate of 2 sccm, the SSA achieves the highest solar thermal conversion efficiency for a solar radiation concentration factor of equal

to or larger than 5. The SSA based on the TiN<sub>x</sub>O<sub>y</sub> layer deposited with the N<sub>2</sub> flow rate of 2 sccm has a solar absorbance of 96.29% and a thermal emittance of 6.11% at the temperature of 400 °C.

## Experimental

The SSA has the structure of SiO<sub>2</sub>-TiO<sub>2</sub>-TiN<sub>x</sub>O<sub>y</sub>-Cu as shown in Figure 1. In this structure, the Cu substrate with high thermal conductivity is utilized as the infrared reflector; the TiN<sub>x</sub>O<sub>y</sub> film serves as the absorbing layer; and the SiO<sub>2</sub> and TiO<sub>2</sub> thin films form a double-layer anti-reflection coating (ARC). The SiO<sub>2</sub>, TiO<sub>2</sub>, and TiN<sub>x</sub>O<sub>y</sub> thin films were deposited by RF magnetron sputtering of 2 inch SiO<sub>2</sub>, TiO<sub>2</sub> and TiN targets (all the targets have a purity > 99.9%), respectively, with a magnetron sputtering system (Denton Desktop Pro.). The chamber was pumped down to 9 × 10<sup>-6</sup> Torr for all the deposition processes. The sputtering was carried out at room temperature, and the sputtering power was maintained at 100 W. The Ar/O<sub>2</sub> gases were used to deposit SiO<sub>2</sub> and TiO<sub>2</sub> thin films with the O<sub>2</sub> flow rate of 1.5 sccm. The TiN<sub>x</sub>O<sub>y</sub> thin films used in the SSA were deposited by Ar/N<sub>2</sub> gases with N<sub>2</sub> flow rate at 2, 5, 8 sccm, respectively. The flow rate of Ar was kept at a constant value of 10 sccm, and the deposition pressure was maintained at around 3 × 10<sup>-3</sup> Torr. The chemical composition of the TiN<sub>x</sub>O<sub>y</sub> thin films were examined with an X-ray photoelectron spectroscopy (XPS) system (ESCALAB 250Xi). The XPS analysis shows that the Ti-O, Ti-N-O, and Ti-N states exist in the TiN<sub>x</sub>O<sub>y</sub> thin films. Figure 2 shows the Ti 2p core level of the two extreme situations of TiN<sub>x</sub>O<sub>y</sub> thin films, one with 0 sccm N<sub>2</sub> flow rate and another with 9 sccm N<sub>2</sub> flow rate (Ar flow rate was maintained at 10 sccm), which were deposited on silicon substrate for the XPS study. Compared to the TiN<sub>x</sub>O<sub>y</sub> thin film with 0 sccm N<sub>2</sub> flow rate, the TiN<sub>x</sub>O<sub>y</sub> thin film with 9 sccm N<sub>2</sub> flow rate has more Ti-O but less Ti-N. This can be explained in the following. The presence of oxygen in the films, which were deposited by sputtering a pure TiN target with Ar/N<sub>2</sub> gases, is due to the existence of residual oxygen in the deposition chamber.<sup>15,16</sup> The presence of oxygen in the TiN<sub>x</sub>O<sub>y</sub> thin films deposited by a sputtering process without purposeful introduction of O<sub>2</sub> gas was also reported previously.<sup>15,17</sup> It was reported that the deposition rate of the

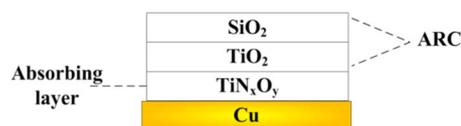
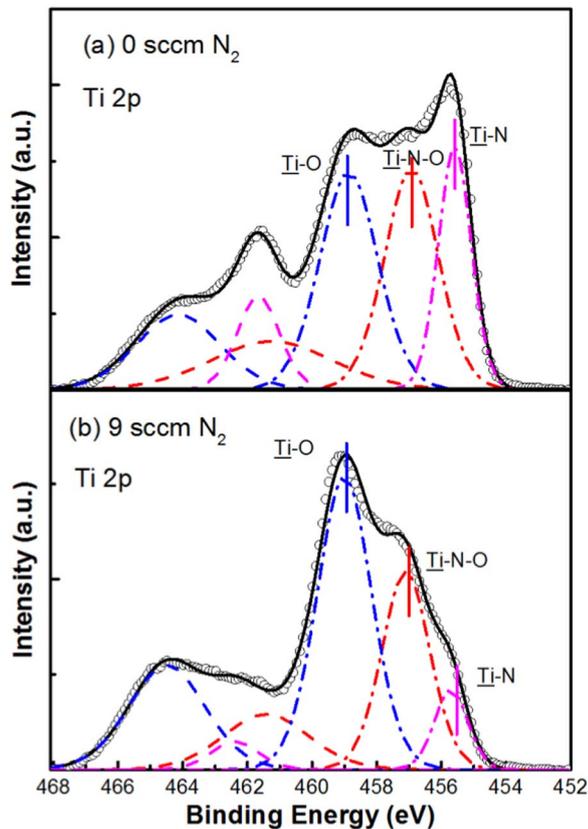


Figure 1. Schematic structure of SSA with SiO<sub>2</sub>-TiO<sub>2</sub>-TiN<sub>x</sub>O<sub>y</sub>-Cu.

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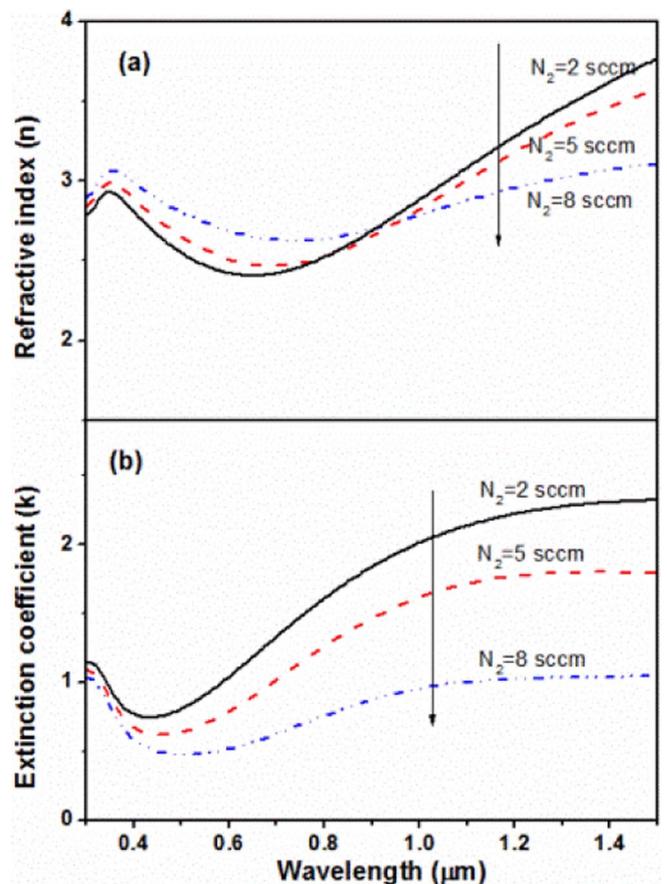


**Figure 2.** XPS analysis of the Ti 2p core level of the two extreme situation of  $\text{TiN}_x\text{O}_y$  thin films on silicon substrate: (a) deposited with 0 sccm  $\text{N}_2$  flow rate; and (b) deposited with 9 sccm  $\text{N}_2$  flow rate. In both situations, the Ar flow rate was maintained at 10 sccm.

$\text{TiN}_x\text{O}_y$  thin films decreases with increase of  $\text{N}_2$  flow rate as a result of the nitriding effect of the TiN target by  $\text{N}_2$ .<sup>15,18</sup> Also note that oxygen has stronger reaction ability than that of nitrogen with Ti and TiN.<sup>6</sup> With the decrease of deposition rate, the residual oxygen has more time to react with the Ti atoms and thus more oxygen is presented in the deposited films. As a result, there is more Ti-O but less Ti-N in the films deposited with a higher  $\text{N}_2$  flow rate. A more detailed XPS examination of the  $\text{TiN}_x\text{O}_y$  thin films is reported elsewhere.<sup>19</sup> The optical properties of  $\text{TiN}_x\text{O}_y$ ,  $\text{TiO}_2$ , and  $\text{SiO}_2$  thin films deposited on silicon substrate were measured with spectroscopic ellipsometry (SE). The SE measurements were conducted with a spectroscopic ellipsometer (Woollam VASE) in the wavelength range of 0.3–1.5  $\mu\text{m}$  at the angle of incidence of  $75^\circ$ . Reflectance of the SSA with  $\text{SiO}_2$ - $\text{TiO}_2$ - $\text{TiN}_x\text{O}_y$  layers deposited on Cu substrate was measured with an UV-Vis-NIR spectrophotometer (PerkinElmer Lambda 950) and a FT-NIR/MIR spectrometer (PerkinElmer Frontier) in the wavelength ranges of 0.3–2.5  $\mu\text{m}$  and 2.5–25  $\mu\text{m}$ , respectively.

## Results and Discussion

To design the layer structure of the SSA, the optical properties of the  $\text{TiN}_x\text{O}_y$  layer and the thicknesses of all the layers of the structure must be optimized. Therefore the optical properties of the  $\text{TiN}_x\text{O}_y$  layer must be accurately determined and their dependence on the deposition conditions should be investigated. In this work, the complex dielectric function ( $\epsilon = \epsilon_r + i\epsilon_m$ , where  $\epsilon_r$  and  $\epsilon_m$  are the real and imaginary parts of the complex dielectric function ( $\epsilon$ ), respectively) of the  $\text{TiN}_x\text{O}_y$  thin film is obtained from the SE analysis based on the Drude-Lorentz dispersion model.<sup>20</sup> As the  $\text{TiN}_x\text{O}_y$  film can be considered as a conductive TiN nanoparticles / oxide composite, the Drude-Lorentz dispersion model takes the contributions of the free electrons, local-



**Figure 3.** Refractive index and extinction coefficient in the wavelength range of 0.3–1.5  $\mu\text{m}$  of the  $\text{TiN}_x\text{O}_y$  thin films deposited with  $\text{N}_2$  flow rates of 2, 5, and 8 sccm, respectively.

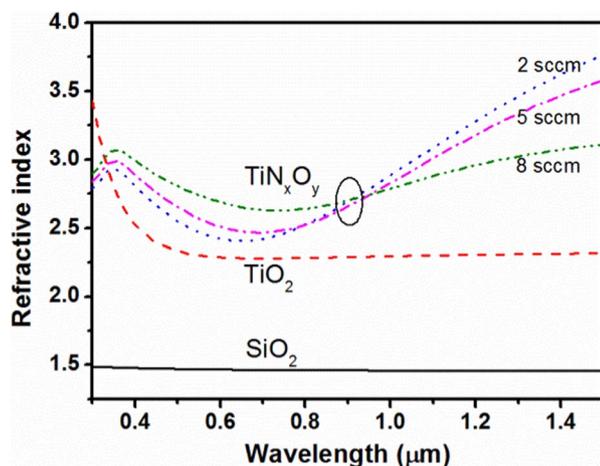
ized surface plasmon resonance (LSPR) and interband transitions into account, which are represented by a Drude term and two Lorentz oscillators, respectively. In the fittings based on the Drude-Lorentz dispersion model to the experimental ellipsometric angles, the mean square error (MSE) was utilized as a criteria.<sup>21</sup> The MSE is smaller than 30 for all the SE fittings. Detail of the methodology of the SE analysis is reported elsewhere.<sup>19</sup>

The complex dielectric functions of the  $\text{TiN}_x\text{O}_y$  thin films deposited with various  $\text{N}_2$  flow rates have been obtained from the SE analysis in the wavelength range of 0.3–1.5  $\mu\text{m}$ . From the complex dielectric functions, the refractive index ( $n$ ) and extinction coefficient ( $k$ ) of the  $\text{TiN}_x\text{O}_y$  thin films are calculated with the following equations:

$$\epsilon_r = n^2 - k^2 \quad [1]$$

$$\epsilon_m = 2nk \quad [2]$$

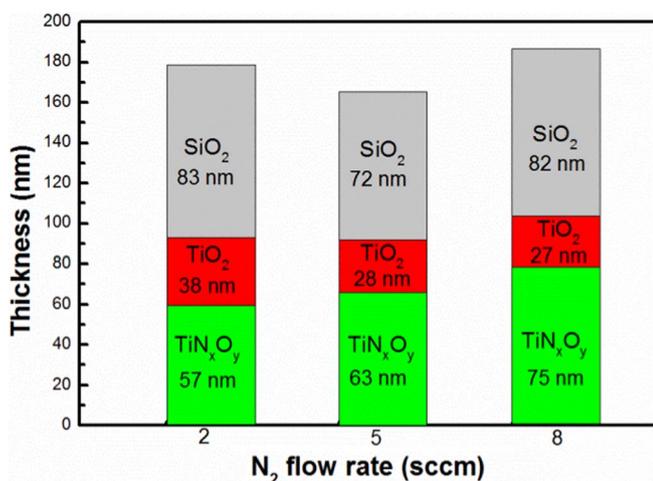
Figure 3 shows the refractive index and extinction coefficient of the  $\text{TiN}_x\text{O}_y$  thin films deposited with  $\text{N}_2$  flow rate of 2, 5, 8 sccm, respectively, in the wavelength range of 0.3–1.5  $\mu\text{m}$ . As shown in Figure 3, with the increasing of  $\text{N}_2$  flow rate, the refractive index of  $\text{TiN}_x\text{O}_y$  increases in the wavelength range of 0.3–0.9  $\mu\text{m}$  and decreases in the wavelength range of 0.9–1.5  $\mu\text{m}$ . The extinction coefficient decreases significantly as the flow rate of  $\text{N}_2$  increases, especially in the infrared range. This indicates that a higher  $\text{N}_2$  flow rate will lead to a lower absorption in the  $\text{TiN}_x\text{O}_y$  layer, which is not suitable to the application of  $\text{TiN}_x\text{O}_y$  as the absorbing layer. Nevertheless, the large dependence of both the refractive index and extinction coefficient of the  $\text{TiN}_x\text{O}_y$  layer on the  $\text{N}_2$  flow rate may provide an opportunity to optimize the SSA layer structure.



**Figure 4.** Refractive indexes of SiO<sub>2</sub>, TiO<sub>2</sub> and TiN<sub>x</sub>O<sub>y</sub> thin films in the wavelength range of 0.3–1.5 μm. The TiN<sub>x</sub>O<sub>y</sub> thin films were deposited with the N<sub>2</sub> flow rates of 2, 5 and 8 sccm, respectively.

In order to achieve a low reflectance in a wide wavelength range (i.e. 0.3–1.5 μm), a graded refractive index profile is usually utilized in a multilayer ARC, which includes low refractive index material (e.g. MgF<sub>2</sub>, SiO<sub>2</sub>, etc.) and high refractive index material (e.g. TiO<sub>2</sub>, Si<sub>3</sub>N<sub>4</sub>, etc.).<sup>22</sup> In this work, SiO<sub>2</sub> and TiO<sub>2</sub> are chosen to form a double-layer ARC on the TiN<sub>x</sub>O<sub>y</sub> layer due to their match in the refractive index profile. As shown in Figure 4, a graded refractive index profile, i.e., low refractive index (SiO<sub>2</sub>) – medium refractive index (TiO<sub>2</sub>) – high refractive index (TiN<sub>x</sub>O<sub>y</sub>) is indeed achieved in the wide spectrum range of 0.34–1.5 μm.

On the other hand, the refractive indexes and extinction coefficients of TiN<sub>x</sub>O<sub>y</sub>, TiO<sub>2</sub> and SiO<sub>2</sub> thin films in the infrared range beyond the wavelength of 1.5 μm, which cannot be determined experimentally, can be obtained by the extension of the results shown in Figure 4 using an appropriate optical dispersion model.<sup>19,23,24</sup> With the knowledge of the optical constants in the spectrum range of 0.3–25 μm, a simulation of the reflectance ( $R(\lambda)$ ) of the SiO<sub>2</sub>-TiO<sub>2</sub>-TiN<sub>x</sub>O<sub>y</sub>-Cu SSA structure can be carried out to minimize the reflectance in the solar radiation range (i.e. 0.3–1.5 μm) and maximize the reflectance in the infrared range (i.e. 1.5–25 μm) by adjusting the thicknesses of the SiO<sub>2</sub>, TiO<sub>2</sub> and TiN<sub>x</sub>O<sub>y</sub> layers. The optimized thicknesses of the three layers for the N<sub>2</sub> flow rates of 2, 5 and 8 sccm are shown in Figure 5. The total thickness of the three layers is around 180 nm for the N<sub>2</sub> flow rates of



**Figure 5.** Optimized thicknesses of the SiO<sub>2</sub>, TiO<sub>2</sub> and TiN<sub>x</sub>O<sub>y</sub> layers used in the SSA for the N<sub>2</sub> flow rates of 2, 5 and 8 sccm, respectively.

2, 5 and 8 sccm. With the increase of N<sub>2</sub> flow rate from 2 to 8 sccm, the thickness of the absorbing layer (i.e. the TiN<sub>x</sub>O<sub>y</sub> layer) increases from 57 to 75 nm due to the decrease of the absorption coefficient of TiN<sub>x</sub>O<sub>y</sub> with increasing N<sub>2</sub> flow rate.

As the SSA layers are deposited on Cu substrate, the transmittance is zero; thus the solar absorbance ( $A_s(\lambda)$ ) at a specific wavelength ( $\lambda$ ) in the solar radiation range can be calculated with Eq. 3.<sup>25</sup> Under the thermal equilibrium condition, the absorbance is equal to the emittance. Therefore the thermal emittance ( $E_{bb}(\lambda)$ ) at a specific wavelength in the infrared radiation range can be obtained with Eq. 4.<sup>26</sup>

$$A_s(\lambda) = 1 - R(\lambda) \quad [3]$$

$$E_{bb}(\lambda) = 1 - R(\lambda) \quad [4]$$

The total solar absorbance ( $A_{ts}$ ) and total thermal emittance ( $E_{tbb}$ ) are given by Eq. 5 and Eq. 6, respectively.<sup>27</sup>

$$A_{ts} = \frac{\int_{\lambda=0.3 \mu\text{m}}^{\lambda=\lambda_c} P_{solar}(\lambda) \times A_s(\lambda) d\lambda}{\int_{\lambda=0.3 \mu\text{m}}^{\lambda=\lambda_c} P_{solar}(\lambda) d\lambda} \quad [5]$$

$$E_{tbb} = \frac{\int_{\lambda=\lambda_c}^{\lambda=25 \mu\text{m}} P_{bb}(\lambda) \times E_{bb}(\lambda) d\lambda}{\int_{\lambda=\lambda_c}^{\lambda=25 \mu\text{m}} P_{bb}(\lambda) d\lambda} \quad [6]$$

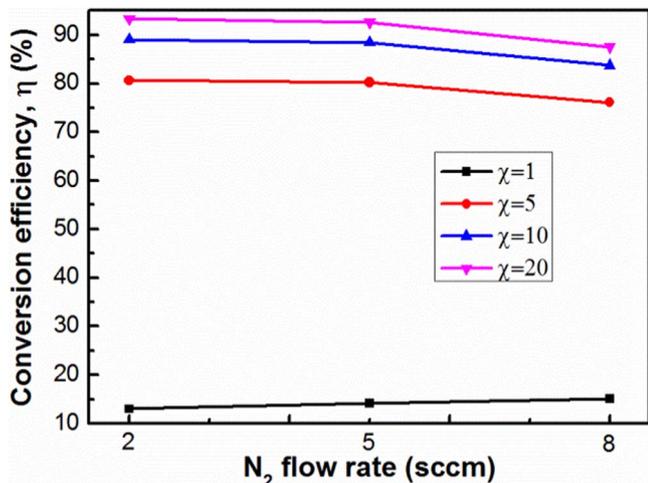
where  $P_{solar}(\lambda)$  is the intensity of the solar radiation (AM 1.5),  $P_{bb}(\lambda)$  is the intensity of black body radiation,  $\lambda_c$  is the transition point.

In order to realize high solar absorbance and low thermal emittance, the reflectance should be low in the solar radiation range and high in the infrared range. The transition point between those two regions is at the cross point of the solar radiation spectrum and the blackbody radiation spectrum. In the mid-temperature SSA application (e.g., 400°C), the blackbody radiation exceeds to the solar radiation at the wavelength of 1.5 μm, as shown in Figure 7. On the other hand, around 90% of solar energy is in the wavelength range of 0.3–1.5 μm.<sup>28</sup> Therefore, the transition point ( $\lambda_c$ ) in this work is set at the wavelength of 1.5 μm.

With the optimized layer thicknesses shown in Figure 5, the total solar absorbance ( $A_{ts}$ ) and total thermal emittance ( $E_{tbb}$ ) of the SSA are calculated with Eq. 5 and Eq. 6, respectively, for N<sub>2</sub> flow rates of 2, 5 and 8 sccm, and the results are shown in Table I. It is found from the table that, with the increase of the N<sub>2</sub> flow rate from 2 to 8 sccm, the  $A_{ts}$  decreases from 97.55% to 91.37%, and the  $E_{tbb}$  also decreases from 7.27% to 6.56%. The change of  $A_{ts}$  with N<sub>2</sub> flow rate can be explained by the change of the chemical stoichiometry of the TiN<sub>x</sub>O<sub>y</sub> thin films. The XPS analysis shows that the ratio of N to O in the TiN<sub>x</sub>O<sub>y</sub> layers decreases from 0.49 to 0.21 with the increase of N<sub>2</sub> from 2 to 8 sccm. This means that the amount of the TiN component in the TiN<sub>x</sub>O<sub>y</sub> layers decreases with N<sub>2</sub> flow rate. The TiN component serves as the absorption center in the TiN<sub>x</sub>O<sub>y</sub> absorbing layers. Therefore, when the N<sub>2</sub> flow rate increases,  $A_{ts}$  decreases with the decrease of the amount of the TiN component. The SSA of the N<sub>2</sub> flow rate of 2 sccm has the largest  $A_{ts}$  (97.55%) due to the largest absorption in the TiN<sub>x</sub>O<sub>y</sub> layer deposited with the N<sub>2</sub> flow rate of 2 sccm. On the other hand, for the N<sub>2</sub> flow rate of 8 sccm, the SSA has the smallest  $E_{tbb}$  (6.56%) at the temperature of 400°C.

**Table I.**  $A_{ts}$  and  $E_{tbb}$  calculated with Eq. 5 and Eq. 6, respectively for the optimized layer thicknesses shown in Figure 5.

	SSA (N <sub>2</sub> = 2 sccm)	SSA (N <sub>2</sub> = 5 sccm)	SSA (N <sub>2</sub> = 8 sccm)
$A_{ts}$	97.55%	96.73%	91.37%
$E_{tbb}$	7.27%	7.10%	6.56%



**Figure 6.** Solar thermal conversion efficiency ( $\eta$ ) at the temperature of 400°C as a function of the concentration factor ( $\chi$ ) for the N<sub>2</sub> flow rates of 2, 5 and 8 sccm.

In order to evaluate the performance of the SSA, the solar thermal conversion efficiency ( $\eta$ ) is calculated with Eq. 7.<sup>29</sup>

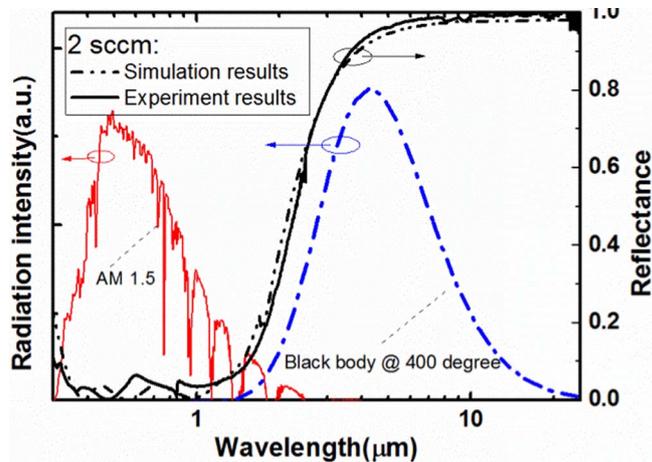
$$\eta = A_{ts} - \frac{E_{tbb}\sigma T^4}{\chi P_{solar}} \quad [7]$$

where  $\sigma$  is the Stefan-Boltzmann constant ( $5.6696 \times 10^{-8} \text{ W} \cdot \text{m}^{-2} \cdot \text{K}^{-4}$ );  $T$  is the absolute temperature;  $\chi$  is the solar concentration factor;  $P_{solar}$  is the total solar radiation intensity. According to Eq. 7, the solar thermal conversion efficiency ( $\eta$ ) depends on the concentration factor ( $\chi$ ). The solar thermal conversion efficiency at the temperature of 400°C as a function of the concentration factor is calculated with Eq. 7 for N<sub>2</sub> flow rates of 2, 5 and 8 sccm, respectively, and the result is shown in Figure 6. When  $\chi$  is one,  $\eta$  is 13.00%, 14.16%, and 15.08%, for the N<sub>2</sub> flow rates of 2, 5 and 8 sccm, respectively, showing a very low  $\eta$ . When  $\chi$  is 5, there is a large increase in  $\eta$ , i.e.,  $\eta$  is 80.64%, 80.22% and 76.11% for the N<sub>2</sub> flow rates of 2, 5 and 8 sccm, respectively. It can be observed from Figure 6 that the N<sub>2</sub> flow rate of 2 sccm provides the highest  $\eta$  for  $\chi$  equal to or larger than 5. In the mid-high temperature SSA applications, the solar energy usually is concentrated on the absorber with parabolic trough with  $\chi$  of larger than 10.<sup>30,31</sup> Therefore, the N<sub>2</sub> flow rate of 2 sccm has been chosen to deposit the TiN<sub>x</sub>O<sub>y</sub> absorbing layer in the fabrication of the mid-high temperature (i.e. 400°C) SSA with the structure of SiO<sub>2</sub>-TiO<sub>2</sub>-TiN<sub>x</sub>O<sub>y</sub>-Cu in this work.

The reflectance of the fabricated SSA with the N<sub>2</sub> flow rate of 2 sccm has been measured in the wavelength range of 0.3–25  $\mu\text{m}$ , and the result is compared with the simulation in Figure 7. As shown in Figure 7, the experiment agrees well with the simulation in the large wavelength range of 0.3–25  $\mu\text{m}$ . The SSA has a low reflectance ( $<0.1$ ) in the wavelength range of 0.3–1.5  $\mu\text{m}$  (i.e. the solar radiation range) and high reflectance in the wavelength range of 1.5–25  $\mu\text{m}$  (i.e. the black body radiation range at 400°C). The measured solar absorbance and thermal emittance at the temperature of 400°C of the fabricated SSA are 96.29% and 6.11%, respectively, which agree well with the calculation discussed early (the calculated solar absorbance and thermal emittance at the temperature of 400°C are 97.55% and 7.27%, respectively).

### Conclusions

By using TiN<sub>x</sub>O<sub>y</sub> as the absorbing layer, a high performance SSA with the structure of SiO<sub>2</sub>-TiO<sub>2</sub>-TiN<sub>x</sub>O<sub>y</sub> coating on Cu substrate has been designed and fabricated. The SiO<sub>2</sub> and TiO<sub>2</sub> layers form a double-layer ARC while the Cu substrate serves as infrared reflector. The TiN<sub>x</sub>O<sub>y</sub> layer was deposited by sputtering of a pure TiN target with



**Figure 7.** Intensity of solar radiation (AM 1.5), intensity of black body radiation at 400°C, and the simulated and experimental reflectance of the SSA with the TiN<sub>x</sub>O<sub>y</sub> layer deposited with the N<sub>2</sub> flow rate of 2 sccm.

Ar and N<sub>2</sub>. Introducing of oxygen in the TiN<sub>x</sub>O<sub>y</sub> layer was realized by utilizing the residual oxygen in the deposition chamber. The N<sub>2</sub> flow rate has a large effect on the optical properties of the TiN<sub>x</sub>O<sub>y</sub> layer; and with the increase of the N<sub>2</sub> flow rate, the extinction coefficient of the TiN<sub>x</sub>O<sub>y</sub> layer decreases significantly. Based on the optical properties of the TiN<sub>x</sub>O<sub>y</sub> layer obtained from the SE analysis, the thicknesses of the SiO<sub>2</sub>, TiO<sub>2</sub> and TiN<sub>x</sub>O<sub>y</sub> layers are optimized to achieve the highest total solar absorbance and the lowest total thermal emittance for a given N<sub>2</sub> flow rate. The N<sub>2</sub> flow rate of 2 sccm provides the highest solar thermal conversion efficiency for the concentration factor equal to or larger than 5. The SSA fabricated with the N<sub>2</sub> flow rate of 2 sccm shows a high solar absorbance of 96.29% and a low thermal emittance of 6.11% at the temperature of 400°C.

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