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<td>Author(s)</td>
<td>Taisne, B.; Jaupart, C.</td>
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Magma degassing and intermittent lava dome growth

B. Taisne and C. Jaupart

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[1] After its 1980 explosive eruption, Mount St Helens developed a lava dome that grew intermittently for several years. Each growth episode was followed by a long repose, suggesting that the magma column above the reservoir was in hydrostatic equilibrium. A mechanism allowing an increasingly thicker dome is proposed. Loading of the crater floor by the dome acts to prevent gas leakage from magma by closing fractures around the volcanic conduit. Fractures get closed down to a depth that increases as the dome grows. Calculations of dome thickness as a function of dome radius are in good agreement with observations. Renewed growth is triggered by the spreading of the dome. Gas retention over a larger depth extent allows smaller magma densities and a taller magma column above the reservoir. According to this model, small domes can in fact promote explosive volcanic conditions and be unstable. This model does not specify what stops magma ascent at the end of a growth phase. At MSH, swelling was short-lived and immediately preceded extrusion after a lengthy repose. There was abundant evidence that the carapace allowed degassing: repose periods were punctuated by small gas explosions and the jetting of gas through open fissures. Thus, swelling was due to the supply of new magma and intermittent dome growth reflected intermittent magma ascent.

[4] There can be no doubt that the various processes invoked above do exist, but they do not explain why successive growth episodes lead to a taller dome. Here, we propose a new model that relies on three observations. One is that the height and diameter of the dome increased continuously. The height to diameter ratio was always smaller than the angle of repose for fractured material and changed with time [Swanson and Holcomb, 1990], showing that the shape of the dome was not determined by static equilibrium conditions. Another important fact is that successive growth episodes were separated by long repose periods with no magma flow in the conduit. A third piece of evidence is the gradual long-term decrease of eruption rate. Dome growth is self-defeating because the weight of the magma/lava column increases until it balances the reservoir pressure. This has been successfully tested in several eruptions [Huppert et al., 1982; Jaupart and Allègre, 1991; Sassiuk et al., 1993]. Renewed magma ascent, therefore, requires either an increase of reservoir pressure or a decrease of the weight of the magma/lava column. At MSH, the overall trend of decreasing eruption rate provides no support for an increase of reservoir pressure. For magma that has nearly constant composition and starting volatile content, changes of the weight of the lava column can only come from modifications of exsolved gas content and/or of dome height. We thus focus on changes of degassing conditions due to loading by the dome and on constraints brought by the dome dimensions. The MSH dome grew to a height of about 250 meters above the vent. Assuming that the density of dome magmas was 1600 kg m$^{-3}$ [Olhoeft et al., 1981], the exit pressure at the vent increased by about 4 MPa. This represents a significant change of pressure, with implications for degassing and the gas content of ascending magma.

1. Introduction

[2] The 1980–1989 eruption of Mount St Helens (MSH) ended with the slow effusion of gas-poor lava accumulating in a dome. In this waning phase, eruption rates were much slower than in the initial explosive phases. Dome growth proceeded through a series of individual episodes typically lasting a few days separated by longer repose periods.

[3] Several models have been put forward to explain intermittent dome growth. Fluctuations of eruption rate can be generated because of kinetically-controlled crystallization during ascent [Melnik and Sparks, 1999]. They may also arise because of the coupling between pressure in the magma reservoir and the eruption rate of crystallizing magma [Melnik and Sparks, 2005] or because of deformation of the conduit walls [Costa et al., 2007]. These models, however, do not account for dome growth and the implied pressure changes. Another possibility is that degassing paths get sealed by mineral precipitation, leading to gas retention and eventually to explosion and dome destruction [Matthews et al., 1997]. Because of dome destruction, the magma/lava column and the reservoir are no longer in hydrostatic equilibrium, which triggers renewed magma flow in the conduit. At MSH, however, there was no significant destruction save for the first two episodes of June and August 1980. In yet another model, swelling of the dome due to the intrusion of new magma or the expansion of magmatic gases proceeds until failure of the dome carapace [Iverson, 1990].

2. Pressure and Gas Content in the Magma Column Below the Vent

[5] The volatile content of MSH magma is about 5 wt% [Rutherford et al., 1985], implying a very large volume fraction of gas at the atmospheric pressure if exsolution and expansion proceed in a closed system. The transition from explosive to effusive eruption regimes reflects the increasing importance of gas leakage from rising magma as the eruption rate decreases [Eichelberger et al., 1986; Jaupart and Allègre, 1991]. Hydrogen isotopic systematics are indeed consistent with open system degassing [Anderson and Fink, 1989].
magma density is to modify the gas content. Not vary appreciably and hence the only way to change position and initial volatile content of the MSH magma did lowering density values in the magma column. The com- rium with the reservoir. A growth increment requires a new phase ended as magma ascent ceased. Thus, growth pro-
ceded until the magma column was in hydrostatic equilib-
rium. Below depth $h_f$, magma leaks all its gas and we assume that its density is constant and equal to $\rho_{mo}$. We start from the reservoir located at depth $H$, where pressure is equal to the lithostatic value plus an overpressure $\Delta P$ (Figure 1). Thus, the magma pressure at $z = h_f$ is:

$$P(h_f) = P_{litho}(H) + \Delta P - \rho_{mo}g(H - h_f)$$

Between $z = h_f$ and $z = 0$, we assume that magma does not lose any gas (closed system behaviour) such that its density depends on gas content $x_g$. The governing equations are:

$$\frac{dP}{dz} = -\rho_{mo}g$$

$$\frac{1}{\rho_m} = \frac{x_g}{\rho_g} + \frac{1-x_g}{\rho_{mo}}$$

$$x_g \approx x[P(h_f)] - x(P)$$

Table 1. Parameters and Physical Properties Used in the Calculation

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<th>Parameters</th>
<th>Properties</th>
<th>Symbol</th>
<th>Value</th>
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<tr>
<td>Density of volatile-free magma</td>
<td>kg $m^{-3}$</td>
<td>$\rho_{mo}$</td>
<td>2600</td>
</tr>
<tr>
<td>Gas density</td>
<td>kg $m^{-3}$</td>
<td>$\rho_g$</td>
<td>calculated as by Jaupart and Allègre [1991]</td>
</tr>
<tr>
<td>Parameter of the solubility law</td>
<td>Pa $^{1/2}$</td>
<td>$s$</td>
<td>$4.11 \times 10^{-5}$</td>
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where $x(P)$ is the solubility of volatile phases in magma. We choose a simple solubility law of the form $x(P) = sP^{1/2}$. Parameters and physical properties are listed in Table 1. We thus calculate $P_o$, the pressure at the volcanic vent, i.e. at the base of the dome. This pressure can also be determined from the dome side, such that $P_o = \rho_d gh_d + P_{atm}$, where $\rho_d$ is the average dome density and $P_{atm}$ is the atmospheric pressure.

Solving the above equations leads to a relationship between the height and radius of the dome. We set the values of $\rho_d$ and $\alpha$ so that $h_f = a r_d$ when $h_f = 0$. This is justified by the observation that the initial domes stalled when they were very thin ($\sim 40$ m). Results are shown in Figure 2 for various values of $\rho_d$ and $\alpha$. This very simple theory predicts a general trend of decreasing height-to-radius ratio which is consistent with the observations. Furthermore, the values of $\rho_d$ and $\alpha$ that are required to fit the data are in good agreement with independent constraints. $\alpha$ must indeed be between values of about 1 and 1.5 and the maximum dome density is about 2600 kg m$^{-3}$, which is that of gas-free MSH magma [Olhoeft et al., 1981]. As shown in Figure 2b, a good fit through the data is achieved for $\alpha = 1$ and $\rho_d = 1750$ kg m$^{-3}$, which is very close to the density of June 1980 dome samples [Olhoeft et al., 1981].

3. Discussion

3.1. Lateral Versus Vertical Gas Escape

One tenet of the model is that gas loss is more effective in the horizontal direction than along the vertical. Two competing effects are involved. On the one hand, if bubbles deform, permeability is enhanced in the direction of bubble elongation, i.e. along the vertical. On the other hand, magma expansion is laterally constrained by the conduit walls, which acts in favor of horizontal permeability development [Llewellin, 2007]. For bubbles that remain spherical, gas percolation occurs first in the horizontal direction [Llewellin, 2007]. Bubble deformation depends on shear stress, which decreases with the eruption rate. Thus, the latter effect is likely to dominate for the small magma velocities that characterize dome-building.

3.2. Long-Term Changes of Dome Properties

Predictions are not in perfect agreement with the data. It would be fruitless to list all the complications that probably come into play and we discuss only one possibility. We have assumed that the average dome density did not change with time and was 1750 kg m$^{-3}$, corresponding to vesicular dacite with about 35% gas bubbles. With time, endogenous growth became more and more important, such that a large fraction of the newly added magma remained within the dome. Such conditions are favorable to gas retention, and it may well be that the average dome density decreased slightly with time. This would explain the slight decrease in density with time.

Figure 2. Calculated dome thickness as a function of dome radius. Horizontal error bars correspond to the minimum and maximum width estimates. Data from Swanson et al. [1987] and Swanson and Holcomb [1990]. (a) Fixed $\alpha = 1$ (such that $h_f = a r_d$) and different average dome densities. (b) Different values of $\alpha$ and fixed dome density ($\rho_d = 1750$ kg m$^{-3}$).

Figure 3. Volume fraction of gas in magma as a function of depth beneath the vent for various values of closure depth $h_f$ (in metres). The vent pressure is determined by the height of the dome (see text). Dome growth has two competing effects on the gas content of magma in the conduit. It inhibits volatile exsolution and gas expansion but also enhances gas retention.
4. Conclusion

[14] Effusion of degassed lava requires gas leakage beneath the volcanic vent, and hence permeable conduit walls. Loading by a dome acts to close fractures at shallow depth in the edifice and to enhance gas retention in the magma. Thus, dome growth can in fact promote explosive volcanic conditions. One expects different behaviours between domes that are free to spread laterally over large distances and domes that are constrained within the confines of small crackers. Our simple model draws attention to the sensitivity of eruption conditions to processes that are active within a few hundred meters of the eruptive vent. It also shows why dome dimensions must be accounted for by eruption models.

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References


Figure 4. Volume fraction of gas in magma at the vent as a function of dome height for various average density values for the dome (in kg m⁻³). Magma does not lose gas between $z = 0$ and $z = h_o$, due to loading by the dome, as specified in the text. Note that the gas content reaches a maximum for a specific dome height.


C. Jaupart and B. Taisne, Dynamique des Fluides Géologiques, Institut de Physique du Globe de Paris, 4, place Jussieu, F-75252 Paris CEDEX 05, France. (taisne@ipgp.jussieu.fr)