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Unique strainmeter observations of Vulcanian explosions, Soufrière Hills Volcano, Montserrat, July 2003


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[1] Five Vulcanian explosions were triggered by collapse of the Soufrière Hills Volcano lava dome in 2003. We report strainmeter data for three explosions, characterized by four stages: a short transition between the onset of disturbance and a pronounced change in strain; a quasi-linear ramp accounting for the majority of strain change; a more gradual continued decline of strain to a minimum value; and a strain recovery phase lasting hours. Remarkable continued decline of strain to a minimum value; and a strain accounting for the majority of strain change; a quasi-steady stage; and a pronounced change in strain; a quasi-steady stage. The duration of vigorous explosion is given by the strain signature, indicating mass fluxes of order 10^17 kg s^{-1}. Conduit pressures released reflect static weight of porous gas-charged magma, and exsolution-generated overpressures of order 10 MPa.


1. Introduction

[2] Most analytical and numerical eruption models address Plinian-type eruptions where vent conditions are quasi-steady. Only a relatively few studies have explored the physics of short-duration Vulcanian eruptions with unsteady conditions, although such events are common at andesitic volcanoes such as Sakurajima and Soufrière Hills Volcano [Ishihara, 1985; Clarke et al., 2002a, 2002b; Druitt et al., 2002; Mason et al., 2006] (see Text S1 in the auxiliary material). The interpretations of such events have been diverse, and there is a need for observations that constrain alternative physical models. Here we use high-precision volumetric strain data and broadband seismic data to constrain the source mechanism, explosion physics and transport dynamics of three well-documented Vulcanian eruptions in 2003 at Soufrière Hills Volcano (SHV), Montserrat, in 2003 [Her et al., 2005].

2. July 2003 Explosions on Montserrat

[3] On 12–13 July 2003, after eight years of eruption at SHV, a major collapse of the lava dome occurred. The collapse followed two years of sustained dome growth and lasted about 18 h. The loss of dome height over the vent was ~500 m, and after collapse the vent was surrounded by crater walls 140–200 m high [Her et al., 2005]. Decompression of magma in the conduit of ~12 MPa caused by loss of overlying dome material triggered a sequence of several Vulcanian explosions. Explosion 1 occurred during the peak of collapse at 03:35 13 July (all times cited are UT), and caused an island-wide fallout of dense juvenile clasts and release of a SO2-rich gas plume [Her et al., 2005]. Explosion 2 occurred at 05:08 during waning stages of the collapse, when the level of ShVC or SHVC seismicity remained high and was dominated by multi-frequency signals from collapse and pyroclastic flow activity [Her et al., 2005].

[4] Three further Vulcanian explosions followed at progressively increasing intervals, after seismicity had returned to background. Explosion 3 occurred at 13:09 13 July, Explosion 4 at 05:15 on 14 July, and Explosion 5 at 05:29 on 15 July [Her et al., 2005]. The explosions were broadly similar in scale, generated lithic and pumice clasts, and were each followed by 2–4 h of ash venting and associated low amplitude tremor. No pyroclastic density currents were detected. Shock waves were reported for several explosions. Washington VAAC reported plumes for Explosion 3 at about 12.2 km altitude, for Explosion 4 about 11.0 km, and for Explosion 5 about 14.3 km high. Strong gas emissions occurred on 16 and 17 July [Edmonds and Herd, 2007], and a small lava dome was extruded 21–28 July, terminating the eruptive episode.

[5] As in 1997 [Druitt et al., 2002], each explosion had an initial, high intensity phase that included the main explosion and peak magma discharge rates, formation of a buoyant plume, and plume ascent to a level of neutral buoyancy in the atmosphere, and a second, waning phase of comparatively negligible discharge, lasting a few hours and characterized by weak venting of ash and gas that generated a low, gradually diminishing bent-over plume. Given the short period of intense discharge compared to explosion plume rise times,
the plumes are modeled as the instantaneous injection of a buoyant thermal of a given initial mass [Morton et al., 1956; Woods and Kienle, 1994; Druitt et al., 2002]. The height \( H_t \) (m) of a thermal is given by \( H_t = 1.89 Q^{0.25} + H_v \), where \( H_v \) is vent height and \( Q \) (J) is the excess thermal mass of injection, taken as the product of \( F \frac{M C}{T} \), where \( F \) is fraction of solids releasing heat to the plume, \( M \) (kg) is mass of solids in plume, \( C \) (J kg\(^{-1}\) K\(^{-1}\)) is specific heat of solids, and \( T \) (K) is initial temperature difference between eruption mixture and ambient air. We assume \( F = 0.8 \), \( C = 1100 \) J kg\(^{-1}\) K\(^{-1}\), \( T = 1123 \) K, and \( H_v = 1000 \) m. We note further that in a humid tropical environment, entrainment of humid air and tropospheric instability can make plume rise exceed the height calculated from the general formula for a given erupted mass, typically by 5–20% [Sparks et al., 1997; Tupper et al., 2009]. To adjust for this observation, we follow the approach of Druitt et al. [2002] and reduce \( H_t \) by 10% prior to calculation.

[6] Given the adjusted plume heights cited previously, the equation yields erupted volumes of 0.42, 0.26 and 0.84 Mm\(^3\) dense rock equivalent (DRE) for Explosions 3–5, respectively, using a reference density of 2600 kg m\(^{-3}\). These volumes are approximate, because in addition to modeling issues, the VAAC height estimates for SHV have an uncertainty of about 1 km that generates a variation range of about 0.62–1.5 times the central volume estimates.

3. Instrumentation Network and Data Description

[7] We present observations of strain and air pressure associated with these explosions, and compare these data with broadband seismic observations (Figures 1 and 2). Data were
obtained from the CALIPSO network [Mattioli et al., 2004], which investigates the dynamics of the Soufrière Hills Volcano (SHV) magmatic system by specialized instruments in four ~200-m-deep boreholes (Figure 1, inset). Three stations (AIRS, TRNT, GERD) were operating for the 13 July explosion, and two (AIRS, TRNT) for 14 and 15 July. Each borehole site, at distances of 5.1 to 9.6 km from the crater (Figure 1a), includes a Sacks-Everson volumetric strainmeter, capable of recording strain changes of about 0.1 nanostrain in periods from DC to about 50 Hz [Linde and Sacks, 1995]. Each strainmeter recorded data at 50 samples per second (sps) with 24-bit ADC’s, and a barograph recorded data at 50 sps with 20-bit resolution. We compare these data with vertical-component broadband seismic data from the surface station MBRY (Figure 1).

The strainmeter responds to ground deformation due to changes in atmospheric pressure, in addition to underground sources, and these effects were removed in order to determine the strain changes from the explosions [Nakao et al., 1989; Linde et al., 2010]. At SHV the explosions typically generated a complicated but intriguing airwave on barographs. The strain shows a rapid decrease in conduit pressure, and the barograph indicates a decrease in air pressure after onset of the eruptive process (Figure 1). For July 2003 events there is no sharp initial positive pressure pulse as observed in 2008–09 explosions, probably because of blocking of the acoustic pulse by crater walls. The remarkable long period ~800 s wave train from these early explosions propagated at about 30 m s⁻¹ as gravity waves for distances >5 km, possibly generated by the rising plume [Kanamori et al., 1994; Lighthill, 1967; Voisin, 1994]. Some shock wave spikes appear in both barometer and strain records during the evacuation process.

The data from AIRS and GERD display good signal/noise ratio (Figure 1). TRNT is located near the coastline and data (not shown) are relatively noisy, but the site is nevertheless important to constrain polarity. For these events the strainmeters at all operating sites expanded (millivolts go negative). AIRS is well beyond the nodal distance for a relatively shallow axisymmetrical source [Linde and Sacks, 1995] (see Text S1), so the source itself has to contract which indicates a decrease in source pressure. The data at AIRS, TRNT and GERD have the same sign and can be satisfied by a relatively shallow (cylindrical) conduit source. The explosion patterns (Figure 1) comprise several stages: 1) a short transition between the onset of disturbance and a pronounced change in strain; 2) a quasi-linear ramp accounting for the majority of strain change; 3) a continued and more gradual decline of strain to a minimum value; and 4) a strain recovery phase lasting hours.

The first of these, Stage 1, is a brief precursor period involving small strain and small seismicity that gradationally precedes the strong explosive release recorded by the onset of large amplitude seismic energy (Figure 2). This precursor period is interpreted to represent initial progressive fracturing as the pressure beneath the conduit plug reaches a failure threshold. In laboratory experiments on samples subjected to a sudden pressure drop [Spieler et al., 2004] tensile failure initiates a fragmentation wave that propagates through the sample. We suggest that Stage 1 occurs when the pressurized gas, accumulating in glassy foam or as a discrete pocket under a strong plug, reaches the tensile strength of the cap [Alidibirov and Dingwell, 1996; Spieler et al., 2004]. Stage 2 denotes a rapid source contraction that we correlate with fragmentation and vigorous explosive evacuation of magma in the conduit. The stage 2 pattern in the raw time series is interrupted by the superposition of very-long-period waves in both barometer and strainmeter signals, but we minimize this influence by applying corrections for air pressure. We propose that most of the ejecta is expelled during the nearly linear strain ramp, and that the subsequent gradual further decline of strain to a minimum value in Stage 3 reflects superposed effects of discharge of ash and degassing at the top of the magma column, and rise of viscous magma in the conduit. Stage 4 continues beyond the strain-minimum value and represents a gradual several hour strain recovery that reflects magma ascent in the conduit and exsolution.

4. Strain Analysis and Other Results

The conduit dimensions are constrained by spine dimensions, magma ascent rates and volume extrusion rates [Voight et al., 1999; Melnik and Sparks, 2002], and we assume for analysis a conduit diameter of 30 m. A transition from a cylindrical conduit to an underlying dike has been proposed for SHV [Mattioli et al., 1998; Costa et al., 2007; Hautmann et al., 2009; Linde et al., 2010], and for analysis we assume a nominal transition depth roughly 2 km beneath the vent at ~700 m a.s.l. There is no suggestion in our strain data that a dike was depressurized, but this fact provides a useful constraint.

The conduit drawdown depth is estimated from erupted DRE volume adjusted for porosity, divided by conduit area. For simplicity we assume the porosity is constant with depth, and we use a average pumice density of ~1500 kg m⁻³, based on data from both the 1997 and 2003 explosions [Clarke et al., 2002b; Edmonds et al., 2006]. We first consider eruption volume. The plume heights suggest erupted volumes of 0.42, 0.26 and 0.84 Mm³ DRE for Explosions 3–5, implying for a 30-m conduit, drawdowns of 1.0, 0.65 and 2.1 km, respectively. The 2.1 km estimate matches within standard error our proposed nominal depth to top of dike.
We also propose another volume-estimation method, based on strain data. We assume eruption volume is proportional to the relative sizes of strain-ramps, with strain-ramp ratios of 0.4, 0.6, and 1.0 for Explosions 3–5, respectively. The procedure is based on the interpretation that the strain-ramps correspond to vigorous evacuations of the conduit, and that other variables such as conduit overpressure are similar. The maximum erupted volume, conditional on a dike top at 2.0 km and 30-m conduit, is 0.81 Mm$^3$ DRE, applied to Explosion 5. Multiplying by strain ratio yields 0.32 and 0.49 Mm$^3$ DRE, for Explosions 3 and 4. Thus the corresponding drawbacks for the three explosions are 0.79, 1.2, and 2.0 km, respectively. We accept the fact of volume uncertainty as previously expressed, and include its consideration in analysis.

Next, we use our strain-ramp amplitudes to estimate average conduit wall pressures released in the explosions. We use an axisymmetrical elastic finite element model (13,000 elements with quadratic shape functions) (see auxiliary material) and a cylindrical conduit of 30 m diameter, applying a magmastic load for a nominal average density of 1500 kg m$^{-3}$ and a superposed uniform overpressure (pressure exceeding magmastic), over the lengths of conduit specified above. We have developed several sub-models, e.g. using homogeneous media for one case, and using inhomogeneous media in another, based on the average one-dimensional seismic velocity model (see auxiliary material), with velocity/modulus increasing with depth, obtained in the SEA-CALIPSO tomography experiment [Shalev et al., 2010]. We assume Young’s modulus $E = 5$ GPa for the homogeneous case; this is the average local modulus value calculated for the strainmeter sites. Similar values have been justified for modeling work at other volcano sites [e.g., Houlé and Montagner, 2007; Cayol and Cornet, 1998], implying a fractured and/or hydrothermally altered rock mass. In all cases we assume a Poisson’s ratio of 0.25. For inhomogeneous media, plume height-based volumes, and Explosions 3, 4, 5, the air pressure-corrected strain-ramps of Stage 2 (8, 12, 20 nano-strain) yield average overpressures of about 15, 50, and 25 MPa, respectively (rounded-off values). For the strain ratio-based volumes, the overpressures are about 25 MPa for the three cases.

The general consistency of average pressures is probably more significant than results for any specific explosion. Further, this result is dependent on modulus: assuming $E = 4$ GPa reduces overpressures $\sim$30%, and $E = 3$ GPa yields overpressures of order 10 MPa. The latter modulus value is considered typical for volcanic areas by Lisowski [2007] (see auxiliary material). For Explosion 3, we can also calculate the ratio of measured strain changes at AIRS and GERD sites (Figure 1) as about 4 (GERD data are not available for the other explosions). The corresponding modeled values are $\sim$5. This result might be explained by modulus anomalies recognized in the 3D tomography [Shalev et al., 2010], not yet considered in our modeling.

To estimate mass evacuation rates, we divide the erupted mass by the corresponding durations of the strain-ramp in stage 2, giving $1.2 \times 10^5, 0.78 \times 10^5, \text{and } 2.1 \times 10^5$ kg s$^{-1}$ for Explosions 3–5, respectively, using plume height-based volumes. For the strain ratio-based volumes, the estimated average rates are $0.57 \times 10^5, 0.91 \times 10^5, 1.3 \times 10^5$ kg s$^{-1}$. The order of magnitude of average mass flux is $10^7$ kg s$^{-1}$, similar to the estimate of Druitt et al. [2002] for the 1997 explosions.

In comparison, model-based estimates of Clarke et al. [2002a, 2002b], Melnik and Sparks [2002] yielded peak discharges of $3-7 \times 10^7$ kg s$^{-1}$.

The volumes erupted in these explosions are necessarily uncertain, with our estimated ranges for Explosions 3–5 being 0.32–0.42 $\times 10^5$, 0.26–0.49 $\times 10^5$, and 0.81–0.84 $\times 10^5$ m$^3$, respectively. These values are similar to the average discharge of $0.3 \times 10^5$ m$^3$ for 88 explosions at SHV in 1997 [Druitt et al., 2002], which occurred at intervals of 2.5–63 h, with a mean of 10 h. However in 2003, the intervals between successive explosions increased: 90 min, to 480 min, to 960 min, to 1450 min, between Explosions 1 and 5 [Herd et al., 2005]. The volumes erupted in the first two explosions are unknown because they occurred during the dome collapse. For refilling the conduit after Explosions 3 and 4, the data suggest average conduit recharge fluxes of $5.6–7.3$ m$^3$ s$^{-1}$, and $3.0–5.6$ m$^3$ s$^{-1}$, respectively. These values are roughly consistent with the enhanced explosion potential noted for fluxes exceeding $\sim$5 m$^3$ s$^{-1}$ [Sparks et al., 1998]. The 2003 explosions appear to be close to the end-member case of zero volatile mass transfer, where magma rising in a conduit acquires a porosity and overpressure structure that evolves to critical conditions to trigger an explosion [Melnik and Sparks, 2002; de’ Michieli Vitturi et al., 2008]. But the decrease noted in recharge flux rates suggests a trend toward more effective magma degassing from the conduit that could explain the delay in meeting the explosion criterion, as well as the lack of significant explosions after 15 July.

### 5. Discussion

Our calculations underscore the importance of overpressures in Vulcanian eruptions, and our modeling suggests average overpressures of about ten to possibly a few tens of MPa. These preliminary results are approximate because of (a) unavoidable eruptive volume uncertainty; (b) uncertainty regarding precise conduit geometry and possible variations with depth; (c) uncertain magma bulk density with depth [de’ Michieli Vitturi et al., 2008]; (d) a simplified pressure distribution in comparison to conduit flow models [Melnik and Sparks, 2002; de’ Michieli Vitturi et al., 2008], and (e) the fact that rock modulus and its distribution is only crudely understood and modeled. Nevertheless, the suggestion of pressurized gas-charged magma at high levels in the conduit immediately prior to the explosion supports the observations of exit velocities $>100$ m/s only a few seconds after Vulcanian explosions began [Druitt et al., 2002; Clarke et al., 2002a, 2002b]. In comparison, Robertson et al. [1998] reported a pressure of 27.5 MPa, based on modeling of ballistic clasts in the 17 September 1996 explosion. This result is similar to some of ours, although we consider our larger estimates as uncomfortably high and emphasize that a reduced modulus can yield overpressures of order 10 MPa. Overpressures estimated by conduit models for magma properties equivalent to those for SHV do not typically exceed 15–20 MPa [de’ Michieli Vitturi et al., 2008; Melnik and Sparks, 2002].

If the explosion threshold is conduit overpressure, then this should in principle be detectable by dilatometer strain. Inspection of the strain records over a three-day period encompassing the explosions shows no clear indications of overpressure buildup. The multiple-day records are, however, uncorrected for tides, ocean loading, or meteorological effects. On timescales of about 1 h we see no evident pressure
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References


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